

Lecture 14

We saw that when  $n \propto n_q$  multiple occupancy of orbitals is possible and quantum statistics must be taken into account

"Degenerate Fermi gases" are ideal gases of fermions under those conditions (here, meaning of degenerate is different than usual meaning in QM where states have same energy)

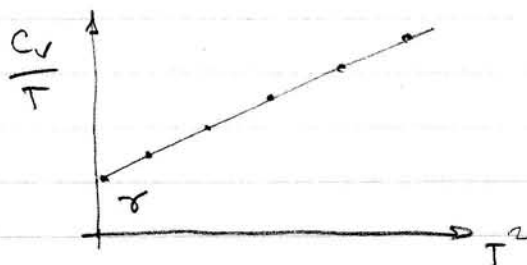
Ex: conduction  $e^-$  in metals - outer  $e^-$  free from ions, conduct  
 white dwarf - atoms are ionized  $e^-$  in gas  
 neutron star - gas of neutrons  
 nuclear matter - gas of protons + neutrons  
 liquid  $^3\text{He}$  - fermions too, but ideal gas approximation not valid

We'll treat conduction  $e^-$  in metals

Recall heat capacity in metals

$$C_V = \gamma T + AT^3$$

$e^- \downarrow \quad \uparrow \text{phonons}$



We'll be able to derive linear dependence of  $C_V$  on  $T$

Recall  $n_q = \left( \frac{2\pi m k_B T}{h^2} \right)^{3/2}$

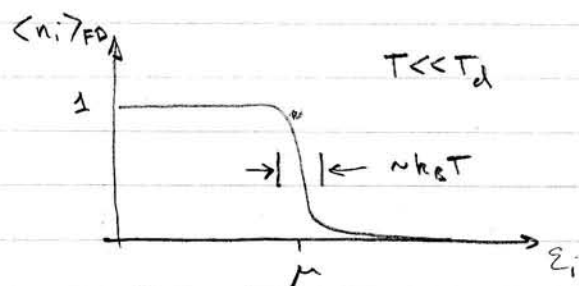
Define degeneracy temperature  $T_d$  when  $n = n_q$ :

$$T_d = \frac{h^2}{2\pi m k_B} n^{2/3}$$

$$\begin{array}{ll} \text{when } T < T_d & n > n_q \\ T > T_d & n < n_q \end{array}$$

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For typical metals  $T_d \sim 50,000 - 100,000 \text{ K}$   
 so  $n \gg n_q$  at room  $T = 300 \text{ K}$ . Metals are degenerate



Occupancy  $\langle n_i \rangle_{FD}$  is very close to a step function ( $T=0$ )  
 (We will use this to approximate solution)

Recall  $N = \sum_i \langle n_i \rangle_{FD} = \sum_i \frac{1}{e^{\beta(\epsilon_i - \mu)} + 1}$  determines  $\mu$

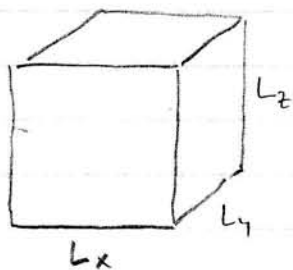
Also  $U = \sum_i \langle \epsilon_i \rangle = \sum_i \langle n_i \rangle_{FD} \epsilon_i = \sum_i \frac{\epsilon_i}{e^{\beta(\epsilon_i - \mu)} + 1}$

From that  $C_V = \left( \frac{\partial U}{\partial T} \right)_V$

At room  $T = 300 \text{ K}$ , can consider spectrum of orbitals to be continuous so  $\sum \rightarrow \int$

$$\left. \begin{aligned} N &= \int_0^\infty d\epsilon D(\epsilon) \langle n(\epsilon) \rangle \\ U &= \int_0^\infty d\epsilon \epsilon D(\epsilon) \langle n(\epsilon) \rangle \end{aligned} \right\} \begin{aligned} D(\epsilon) &= \text{density of states} \\ D(\epsilon) d\epsilon &= \# \text{ of orbitals with energy between } \epsilon, \epsilon + d\epsilon \end{aligned}$$

To get density of states, use same approach as before



$e^-$  in box of volume  $L_x L_y L_z = V$

$e^-$  wavefunctions quantized

$$k_x = \frac{n_x \pi}{L_x}, \quad n_x = 1, 2, 3, \dots$$

same for  $y, z$

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so each state takes up volume in  $k$ -space  $dk_x dk_y dk_z = \frac{\pi^3}{V}$

Energy of state  $\epsilon_k = \frac{\hbar^2 k^2}{2m}$  with  $k = (k_x^2 + k_y^2 + k_z^2)^{1/2}$

# States with energy  $\epsilon, \epsilon + d\epsilon = \# \text{ states in shell } \frac{4\pi}{8} k^2 dk$

$$\therefore D(\epsilon) d\epsilon = 2 \cdot \frac{4\pi \hbar^2}{8} dk \cdot \frac{V}{\pi^3} = \frac{V}{\pi^2} k^2 dk$$

2 possible spins for  $e^-$  :  $\uparrow$  or  $\downarrow$

$$k = \left( \frac{2m\epsilon}{\hbar^2} \right)^{1/2} \quad dk = \left( \frac{2m}{\hbar^2} \right) \frac{d\epsilon}{2\epsilon^{1/2}}$$

$$\therefore \boxed{D(\epsilon) d\epsilon = \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \epsilon^{1/2} d\epsilon}$$

So, we have

$$N = g_0 \int_0^\infty d\epsilon \frac{\epsilon^{1/2}}{e^{\beta(\epsilon - \mu)} + 1} \quad \text{defined } g_0 \equiv \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2}$$

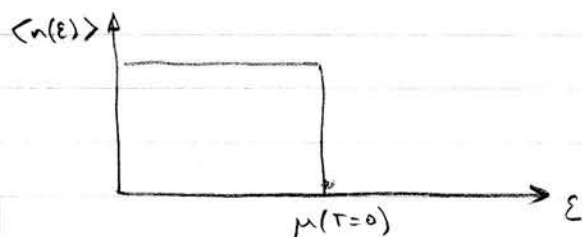
$$U = g_0 \int_0^\infty d\epsilon \frac{\epsilon^{3/2}}{e^{\beta(\epsilon - \mu)} + 1}$$

In general, we cannot solve these integrals analytically. However, at  $T=0$   $\langle n(\epsilon, T=0) \rangle$  is a step function, and integrals are trivial.

Since at room  $T$   $\langle n(\epsilon, T) \rangle$  is close to  $\langle n(\epsilon, 0) \rangle$  we will expand about the  $T=0$  solution to get an approximate answer.

This is called the Sommerfeld expansion.

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First look at  $T=0$ .

$$\begin{aligned} \langle n(\epsilon, 0) \rangle &= 0 & \text{if } \epsilon > \mu(0) \\ &= 1 & \text{if } \epsilon < \mu(0) \end{aligned}$$

$$\begin{aligned} \therefore N &= \int_0^{\infty} \langle n(\epsilon, 0) \rangle D(\epsilon) d\epsilon = \int_0^{\mu(0)} D(\epsilon) d\epsilon \\ &= \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \int_0^{\mu(0)} \sqrt{\epsilon} d\epsilon = \frac{2}{3} \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \mu(0)^{3/2} \end{aligned}$$

$$\boxed{\mu(0) = \frac{\hbar^2}{2m} \left( 3\pi^2 \frac{N}{V} \right)^{2/3} \equiv \epsilon_F}$$

Called the Fermi Energy

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—  $\uparrow\downarrow$  — — —  $\epsilon_F = \mu(0)$  $\uparrow\downarrow$  $\uparrow\downarrow$  $\uparrow\downarrow$ 

:

 $\epsilon_F$  is the highest energy orbital occupied with  $e^-$ Recall that  $e^-$  will fill up orbitals, at  $T=0$  occupancy goes to 0 as soon as  $N$   $e^-$  have been used upNotice  $\mu(0)$  depends on  $N$ , also  $V$  because that affects spacing between energy levels  
 $\mu(0) \sim n^{2/3}$  is intensive

$$\begin{aligned} \text{Now } U(0) &= \int_0^{\infty} \langle n(\epsilon, 0) \rangle D(\epsilon) \epsilon d\epsilon = \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \int_0^{\epsilon_F} \epsilon^{3/2} d\epsilon \\ &= \frac{2}{5} \frac{V}{2\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \epsilon_F^{5/2} = \frac{3}{5} N \epsilon_F \end{aligned}$$

Compare to classical gas with MB statistics

$$U_{MB} = \frac{3}{2} N k_B T, \quad U_{FD} = \frac{3}{5} N \epsilon_F$$

Even at  $T \approx 0$ ,  $U_{FD} \gg 0$  because Pauli exclusion principle requires that  $e^-$  keep filling up energy levels

Define Fermi temperature  $k_B T_F = \epsilon_F$

(Almost the same as degeneracy temp  $T_d$ )

$$k_B T_d = \frac{h^2}{2\pi m} n^{2/3} = 4\pi \frac{\hbar^2}{2m} n^{2/3} \sim (3\pi^2)^{2/3} \frac{\hbar^2}{2m} n^{2/3} = k_B T_F$$

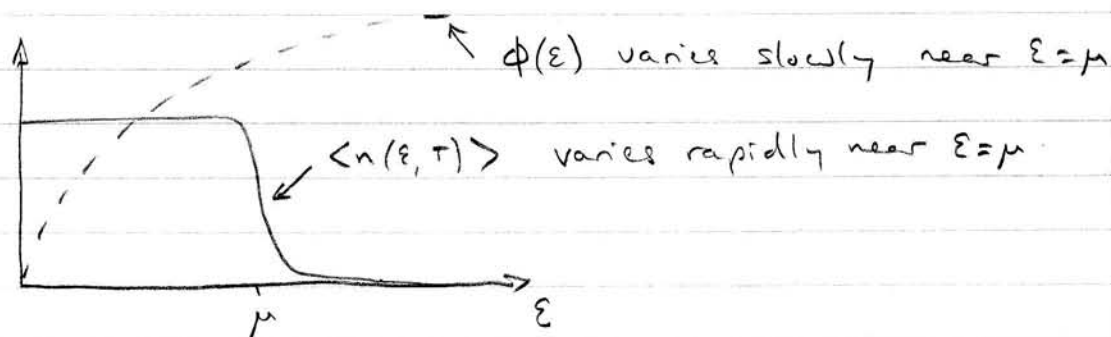
$T_F \sim 50,000 - 100,000 \text{ K}$  in metals )

Now consider  $T > 0$ . Since  $\langle n(\epsilon, T) \rangle$  is close to  $\langle n(\epsilon, 0) \rangle$  at room  $T$ , we'll use the Sommerfeld expansion

We're dealing with integrals of the form

$$I = \int_0^\infty d\epsilon \langle n(\epsilon, T) \rangle \epsilon^n D(\epsilon) \quad \text{For } N, n=0, \quad U, n=1$$

Define  $\phi(\epsilon) \equiv \epsilon^n D(\epsilon)$



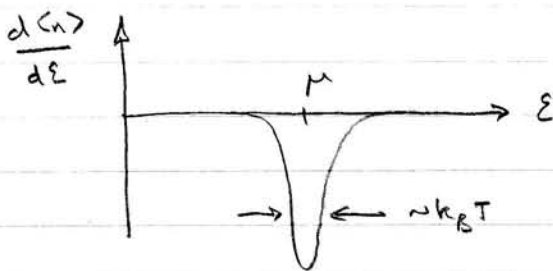
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Sommerfeld Expansion

Define  $\Psi(\varepsilon) \equiv \int_0^\varepsilon \phi(\varepsilon') d\varepsilon'$

Integrate  $I$  by parts

$$\begin{aligned}
 I &= \int_0^\infty \langle n(\varepsilon, T) \rangle \phi(\varepsilon) d\varepsilon \\
 &= \underbrace{\langle n(\varepsilon, T) \rangle}_{n(\infty, T) = 0} \underbrace{\Psi(\varepsilon)}_{\Psi(0) = 0} \Big|_0^\infty - \int_0^\infty d\varepsilon \Psi(\varepsilon) \frac{d\langle n(\varepsilon, T) \rangle}{d\varepsilon} \\
 &= - \int_0^\infty d\varepsilon \Psi(\varepsilon) \frac{d\langle n(\varepsilon, T) \rangle}{d\varepsilon}
 \end{aligned}$$



$\frac{d\langle n \rangle}{d\varepsilon}$  is a very sharp function peaked at  $\varepsilon = \mu$  width of  $\sim k_B T$

$\Psi(\varepsilon)$  varies slowly near  $\varepsilon = \mu$

$\therefore$  Expand  $\Psi(\varepsilon)$  about  $\varepsilon = \mu$

$$\begin{aligned}
 \Psi(\varepsilon) &= \Psi(\mu) + \Psi'(\mu) (\varepsilon - \mu) + \frac{1}{2} \Psi''(\mu) (\varepsilon - \mu)^2 + \dots \\
 &= \sum_{m=0}^{\infty} \frac{1}{m!} \Psi^{(m)}(\mu) (\varepsilon - \mu)^m
 \end{aligned}$$

$$\therefore I = - \sum_{m=0}^{\infty} \frac{1}{m!} \Psi^{(m)}(\mu) \int_0^\infty (\varepsilon - \mu)^m \frac{d\langle n(\varepsilon, T) \rangle}{d\varepsilon} d\varepsilon$$

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$$\frac{d\langle n \rangle}{d\varepsilon} = \frac{-\beta e^{\beta(\varepsilon-\mu)}}{(e^{\beta(\varepsilon-\mu)} + 1)^2}$$

$$\therefore I = \sum_{m=0}^{\infty} \frac{1}{m!} \Psi^{(m)}(\mu) \underbrace{\int_0^{\infty} \frac{\beta e^{\beta(\varepsilon-\mu)}}{(e^{\beta(\varepsilon-\mu)} + 1)^2} (\varepsilon-\mu)^m d\varepsilon}$$

Change variables to  $x \equiv \beta(\varepsilon-\mu)$

$$= \beta^{-m} \int_{-\beta\mu}^{\infty} \frac{e^x}{(e^x + 1)^2} x^m dx$$

Look at lower limit  $\mu \sim \varepsilon_F$  so  $\beta\mu = \mu/k_B T \gg 1$

Also integrand is negligible for  $x < 0$ , so its fine to replace lower limit with  $-\infty$

So integral becomes 
$$I_m = \int_{-\infty}^{\infty} \frac{e^x x^m}{(e^x + 1)^2} dx$$

Note that  $\frac{e^x x^m}{(e^x + 1)^2} = \frac{x^m}{(e^x + 1)(1 + e^{-x})}$  which is an odd function if  $m$  is odd

$$\therefore I_m = 0 \quad \text{for } m \text{ odd}$$

For  $m$  even, we'll only need the two lowest order terms

$$I_0 = \int_{-\infty}^{\infty} \frac{e^x}{(e^x + 1)^2} dx = \left. \frac{-1}{e^x + 1} \right|_{-\infty}^{\infty} = 1$$

$$I_2 = \int_{-\infty}^{\infty} \frac{e^x x^2}{(e^x + 1)^2} dx = 2 \zeta(2) = \frac{\pi^2}{3}$$

↑  
Riemann  
zeta function

(Deriving this takes a bit more algebra)

$$\begin{aligned} \text{So: } \int_0^\infty \langle n(\epsilon, T) \rangle \phi(\epsilon) d\epsilon &= \sum_{\substack{m \\ \text{even}}} \frac{1}{m!} \Psi^{(m)}(\mu) (k_B T)^m I_m \\ &= \Psi(\mu) + (k_B T)^2 \frac{\pi^2}{6} \Psi''(\mu) + \dots \end{aligned}$$

Recall that  $\Psi(\epsilon) = \int_0^\epsilon \phi(\epsilon') d\epsilon'$

$$\therefore \boxed{\int_0^\infty \langle n(\epsilon, T) \rangle \phi(\epsilon) d\epsilon = \int_0^\mu \phi(\epsilon) d\epsilon + \frac{\pi^2}{6} (k_B T)^2 \phi'(\mu) + \dots}$$

the Sommerfeld expansion!

Now consider first integral

$$N = \int_0^\infty \langle n(\epsilon, T) \rangle D(\epsilon) d\epsilon = \int_0^\mu D(\epsilon) d\epsilon + \frac{\pi^2}{6} (k_B T)^2 D'(\mu) + \dots$$

$$N \approx \underbrace{\int_0^{\epsilon_F} D(\epsilon) d\epsilon}_{T=0 \text{ term}} + \underbrace{\int_{\epsilon_F}^\mu D(\epsilon) d\epsilon}_{\approx D(\epsilon_F)(\mu - \epsilon_F)} + \frac{\pi^2}{6} (k_B T)^2 D'(\mu)$$

$$\begin{aligned} T=0 \text{ term} &\approx D(\epsilon_F)(\mu - \epsilon_F) \\ &= N \end{aligned}$$

expect  $\mu(T) \approx \epsilon_F$

$$0 \approx D(\epsilon_F)(\mu - \epsilon_F) + \frac{\pi^2}{6} (k_B T)^2 D'(\epsilon_F)$$

Solve for  $\mu(T)$ :

$$\mu(T) \approx \epsilon_F - \frac{\pi^2}{6} (k_B T)^2 \underbrace{\frac{D'(\epsilon_F)}{D(\epsilon_F)}}_{\frac{1}{2\epsilon_F}} = \epsilon_F - \frac{\pi^2}{12} \left( \frac{k_B T}{\epsilon_F} \right) k_B T$$

this is very small

So  $\mu(T)$  decreases from  $\mu(0) = \epsilon_F$  as  $T$  increases